

Virgo detektor fal grawitacyjnych



Prezentacja na
seminarium Katedry
Radioastronomii

28 X 2010 r.

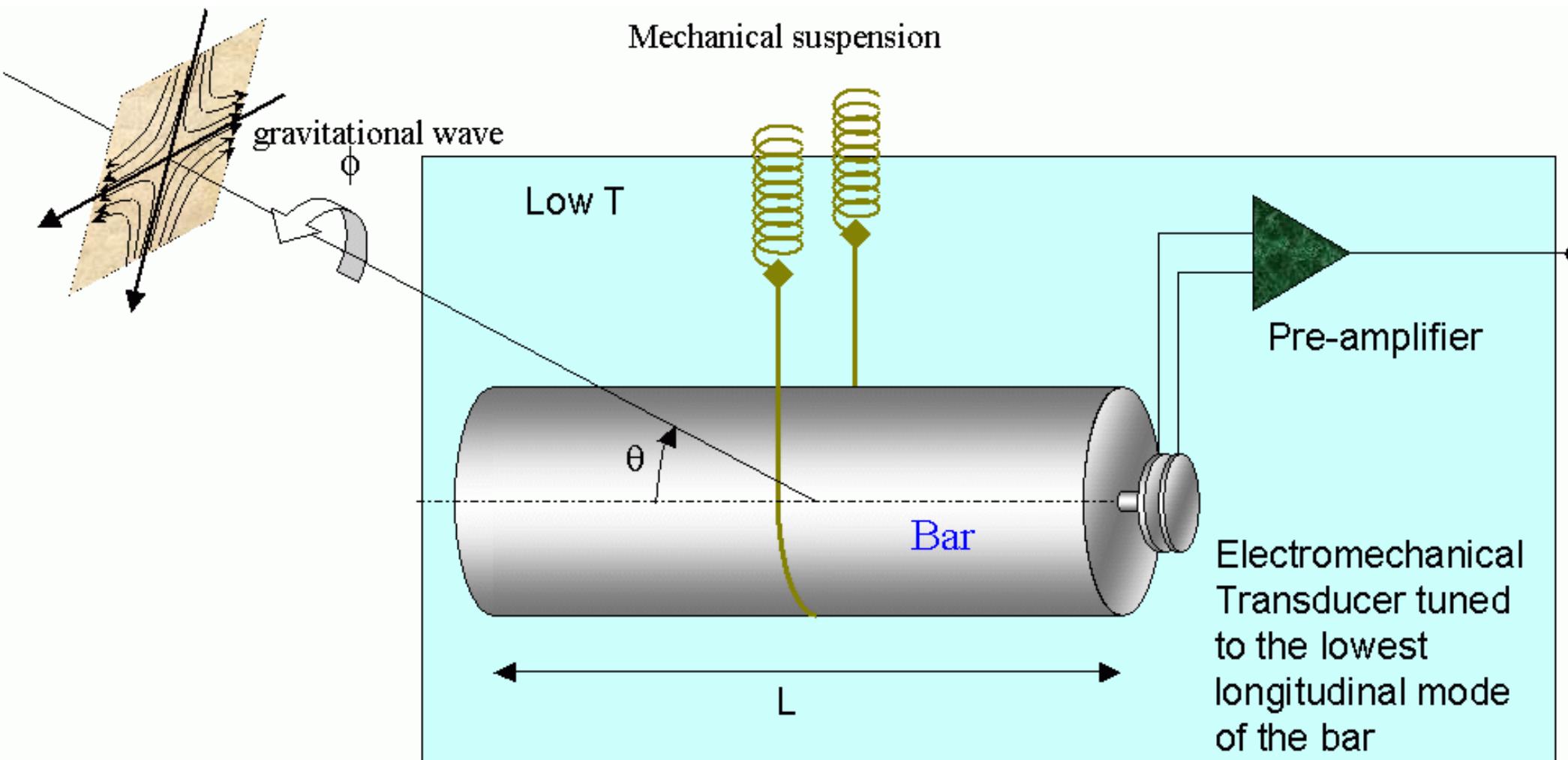
Kazimierz Borkowski

- Detektory fal grawitacyjnych na świecie
Detektory w naszych grantach:
Explorer, Nautilus i Virgo
- Virgo
- Mój dyżur przy Virgo (IX 2010)
- Polska grupa analizy danych (POLGRAW)
Skład osobowy i publikacje
- Źródła
(<http://www.ego-gw.it/virgodescription/indice.html>)

Detektory fal grawitacyjnych



Idea detektora walcowego

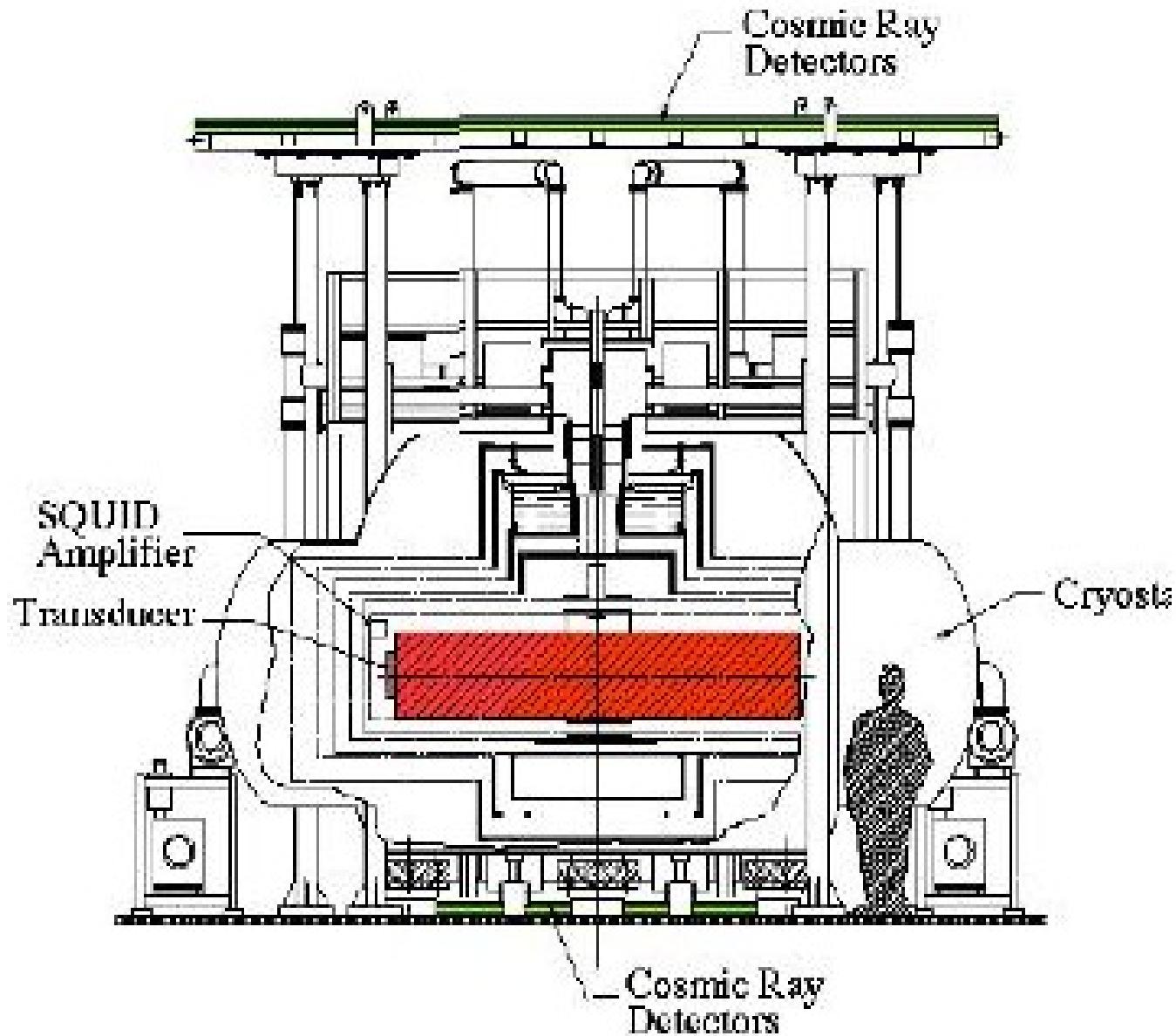


Explorer





Nautilus

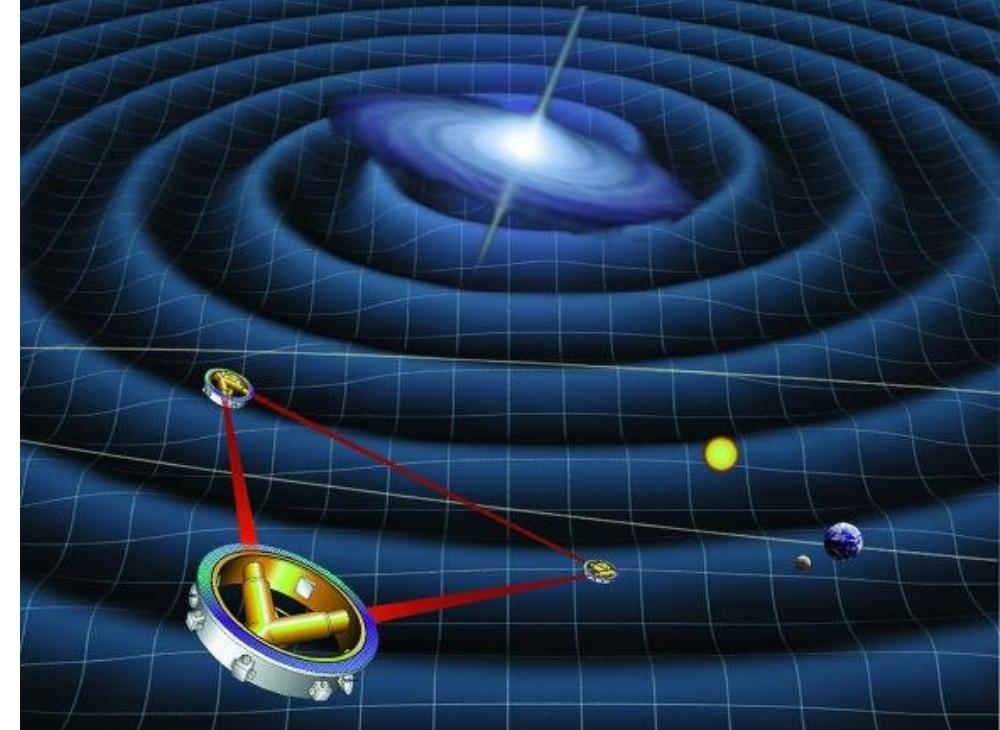




MiniGRAIL
Leiden



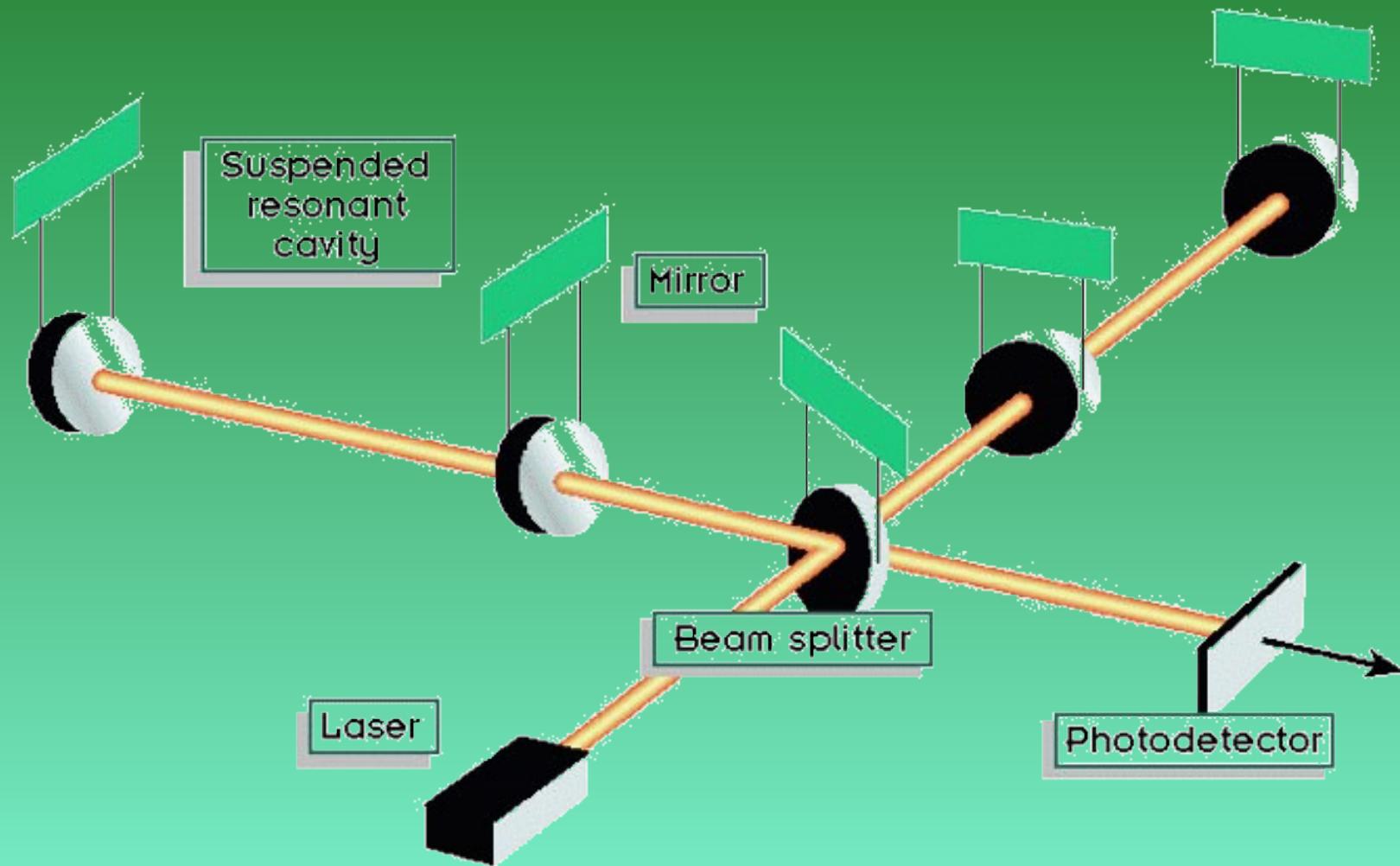
LIGO – Washington (left), Luisiana (right)

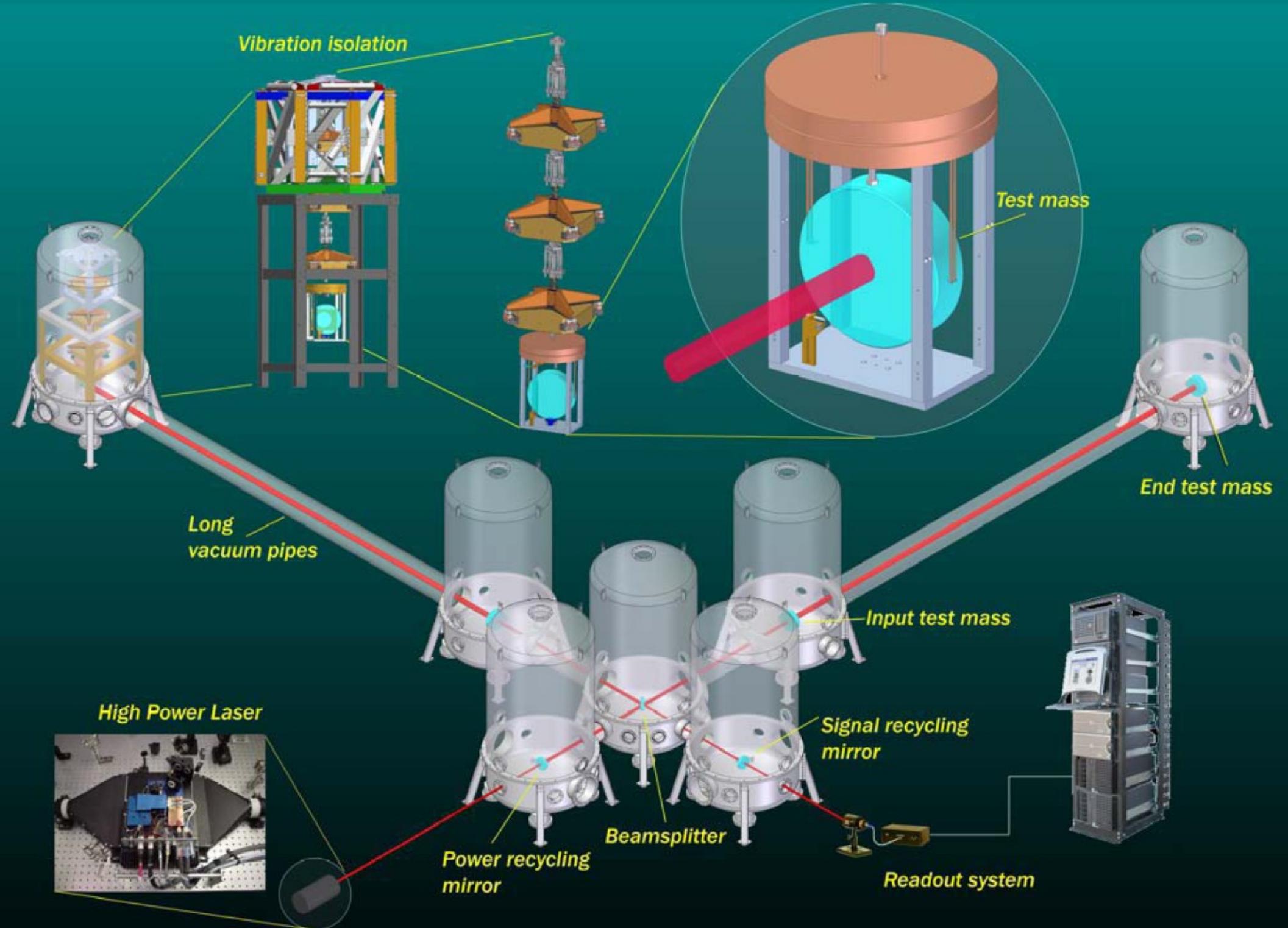


LISA - NASA/ESA

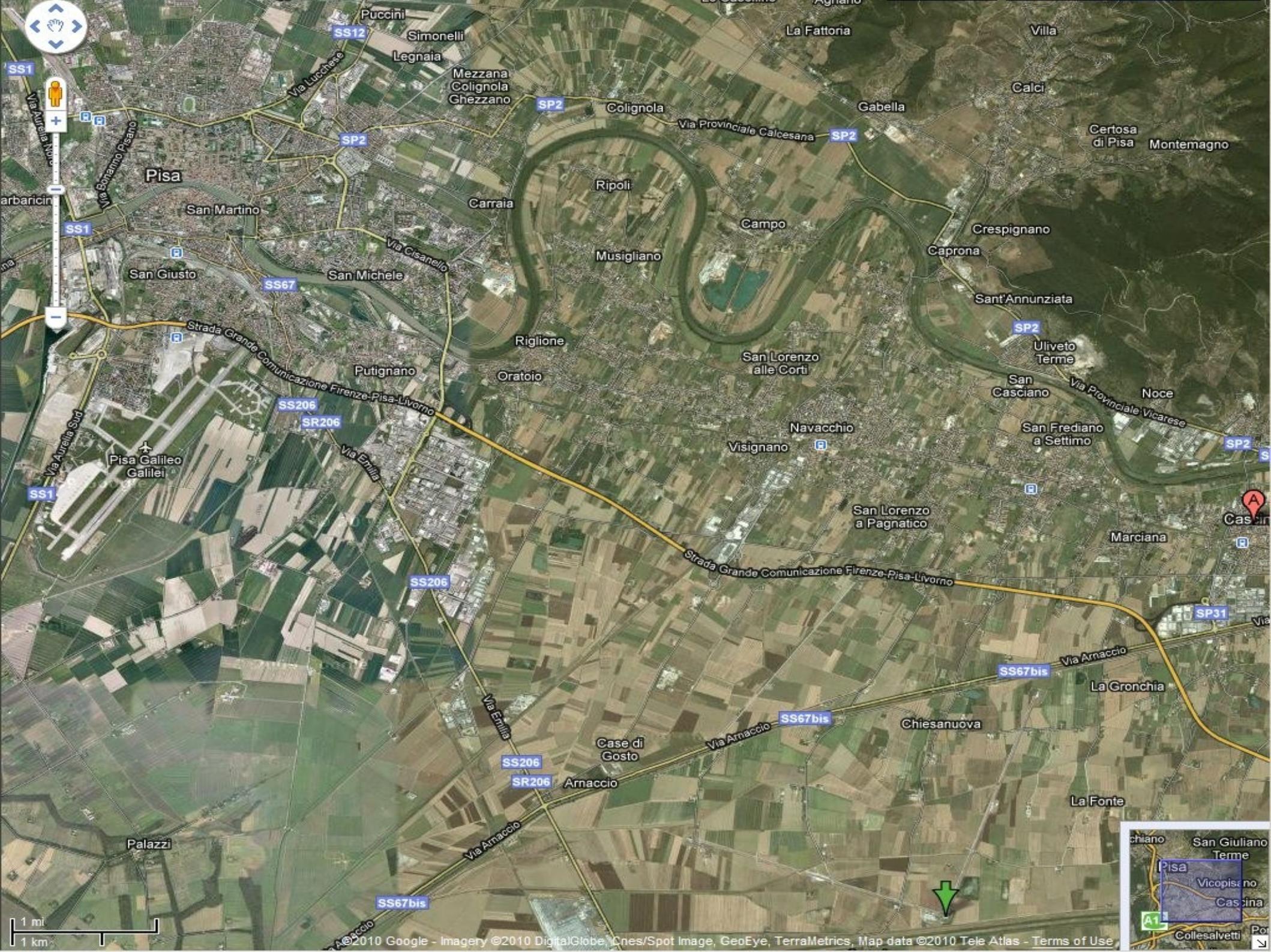
Leiden

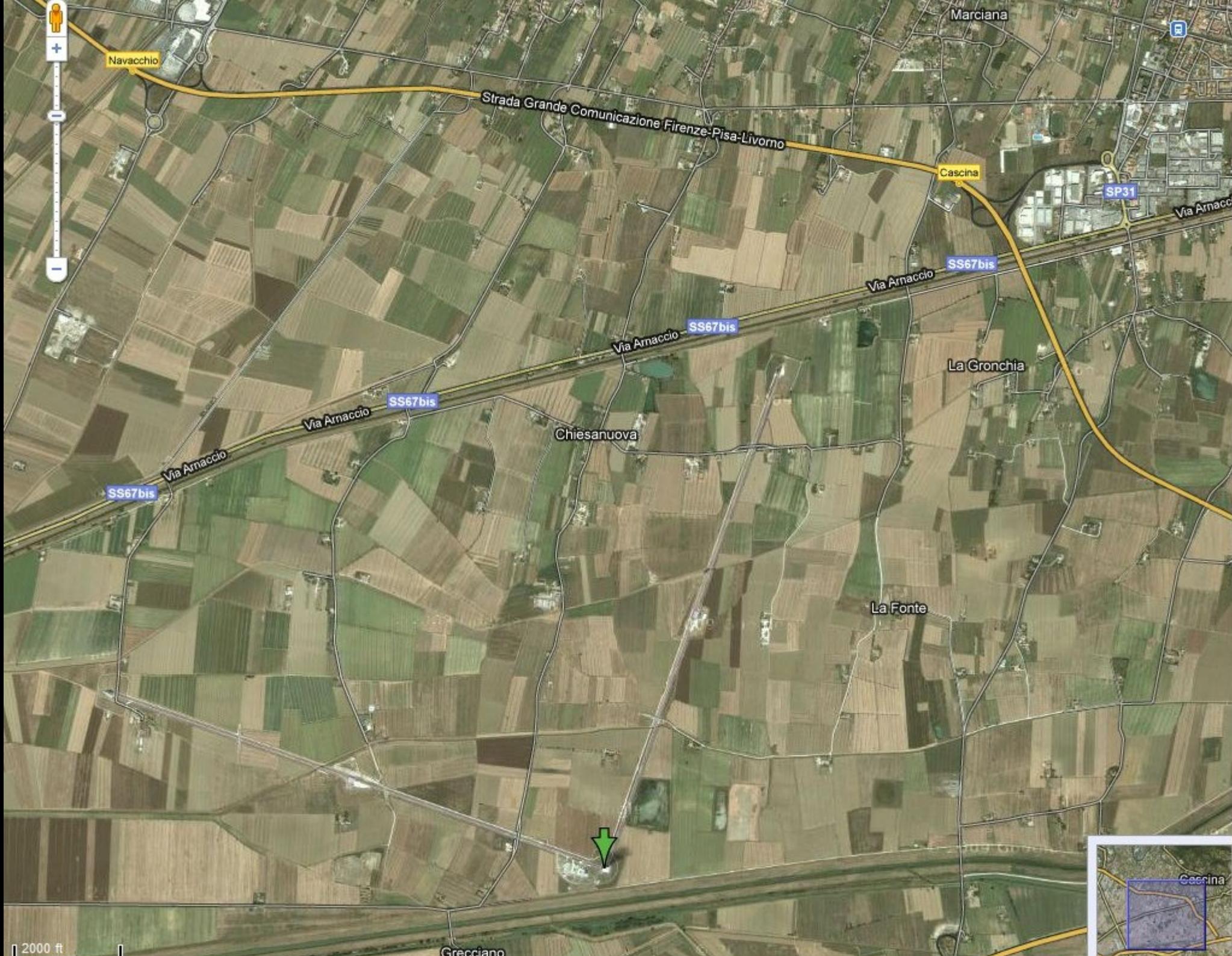
Idea detektora laserowego









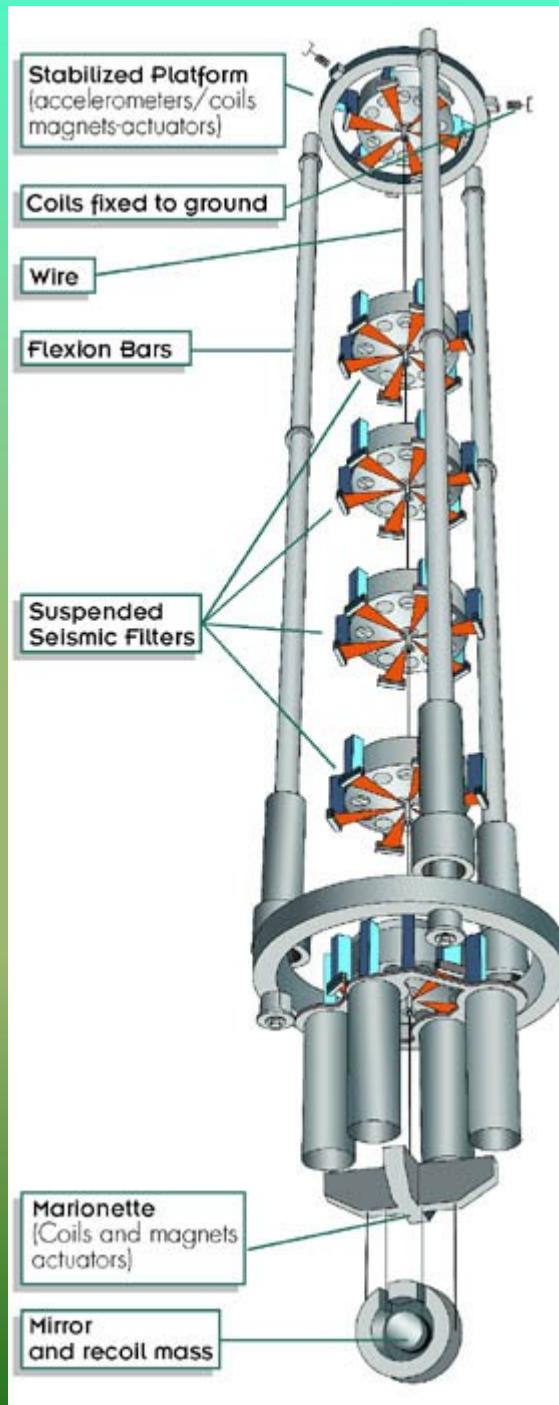


2000 ft



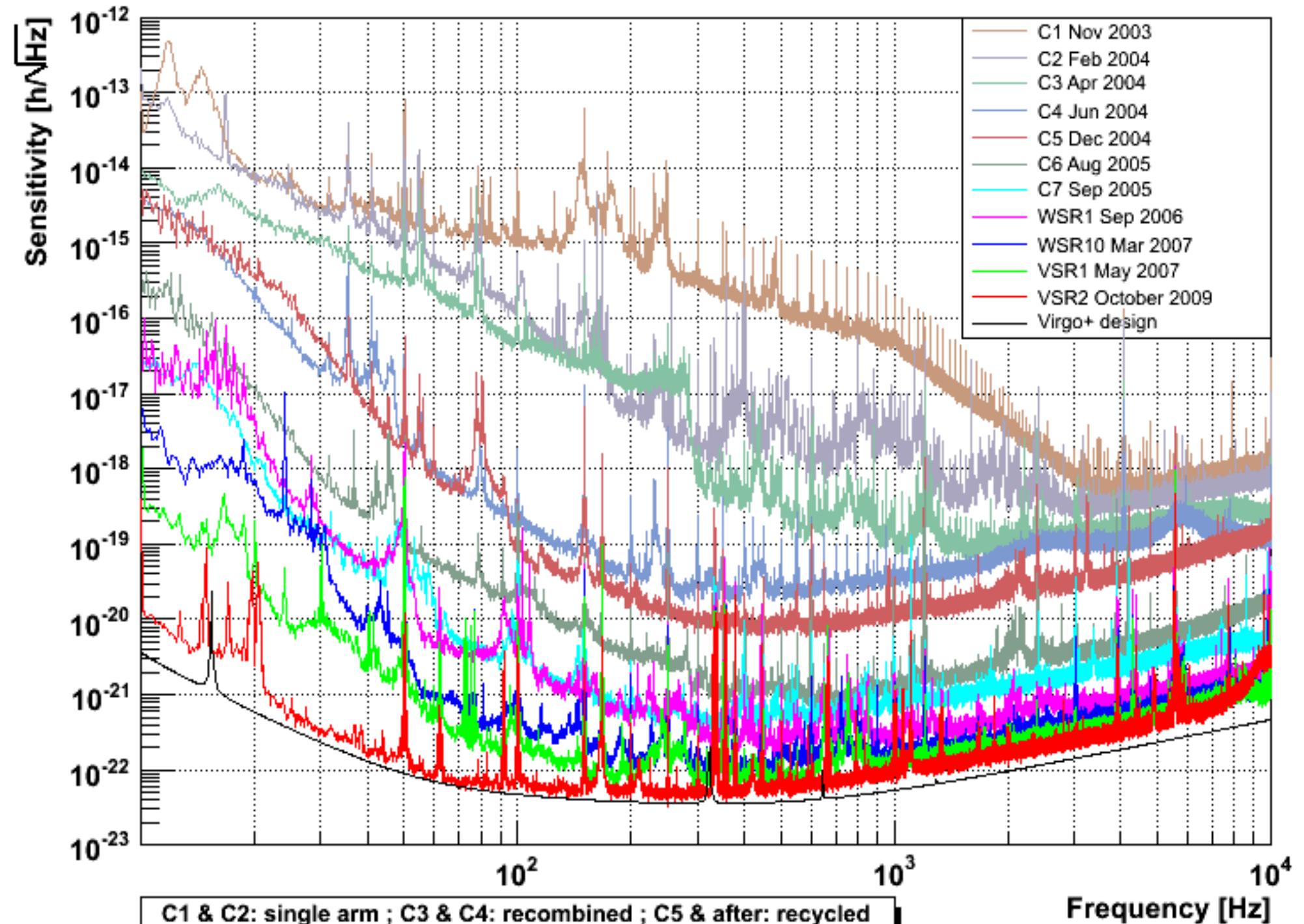


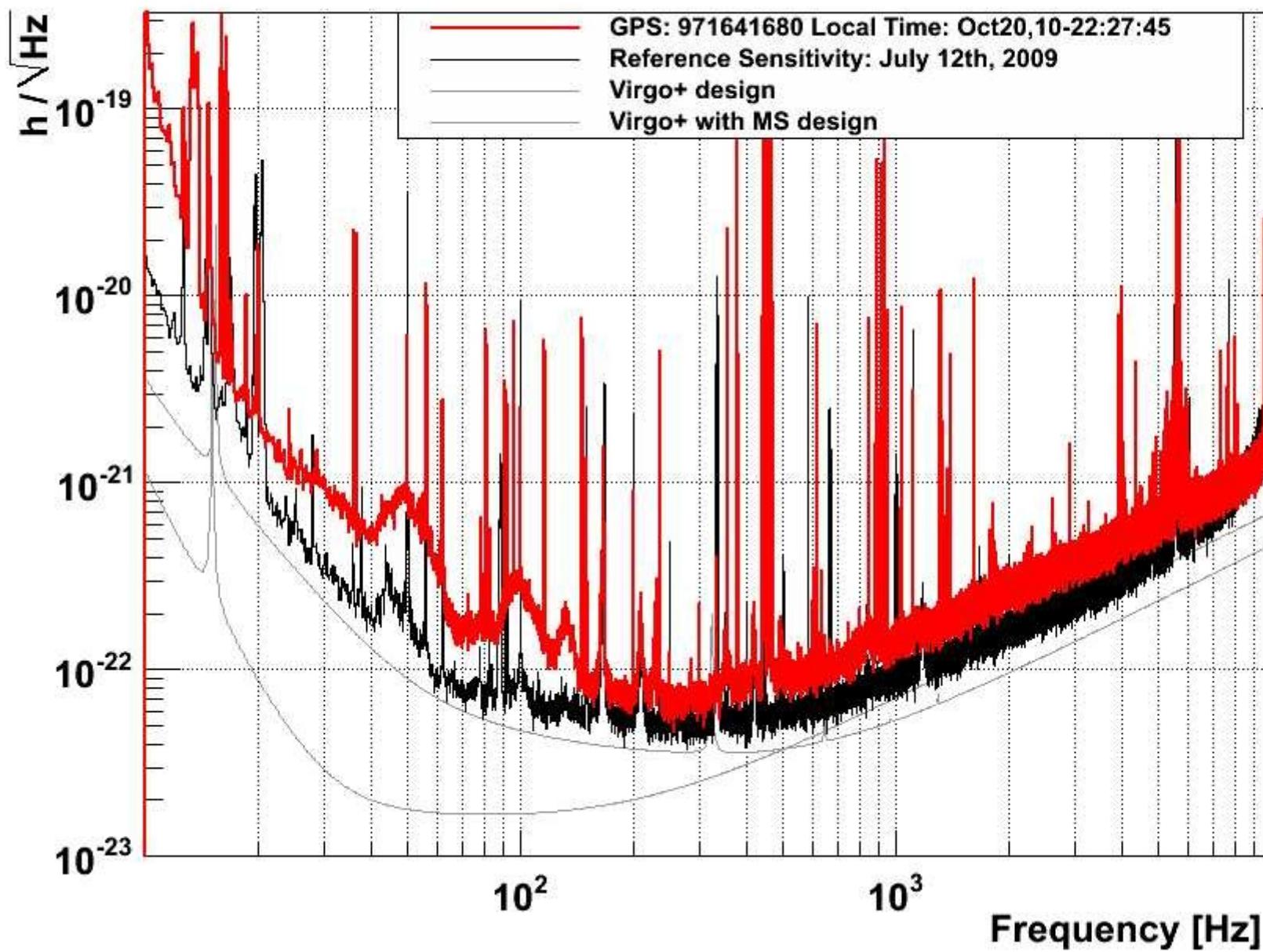


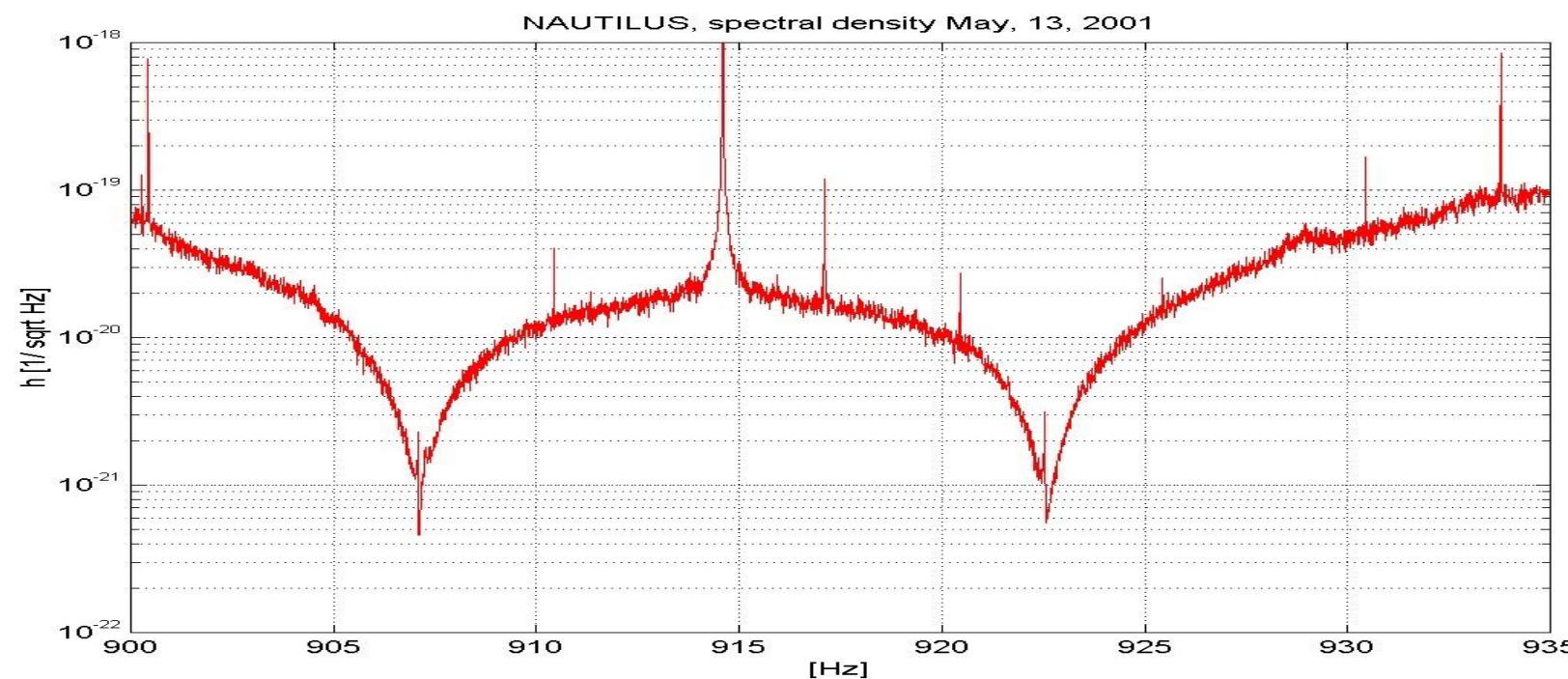
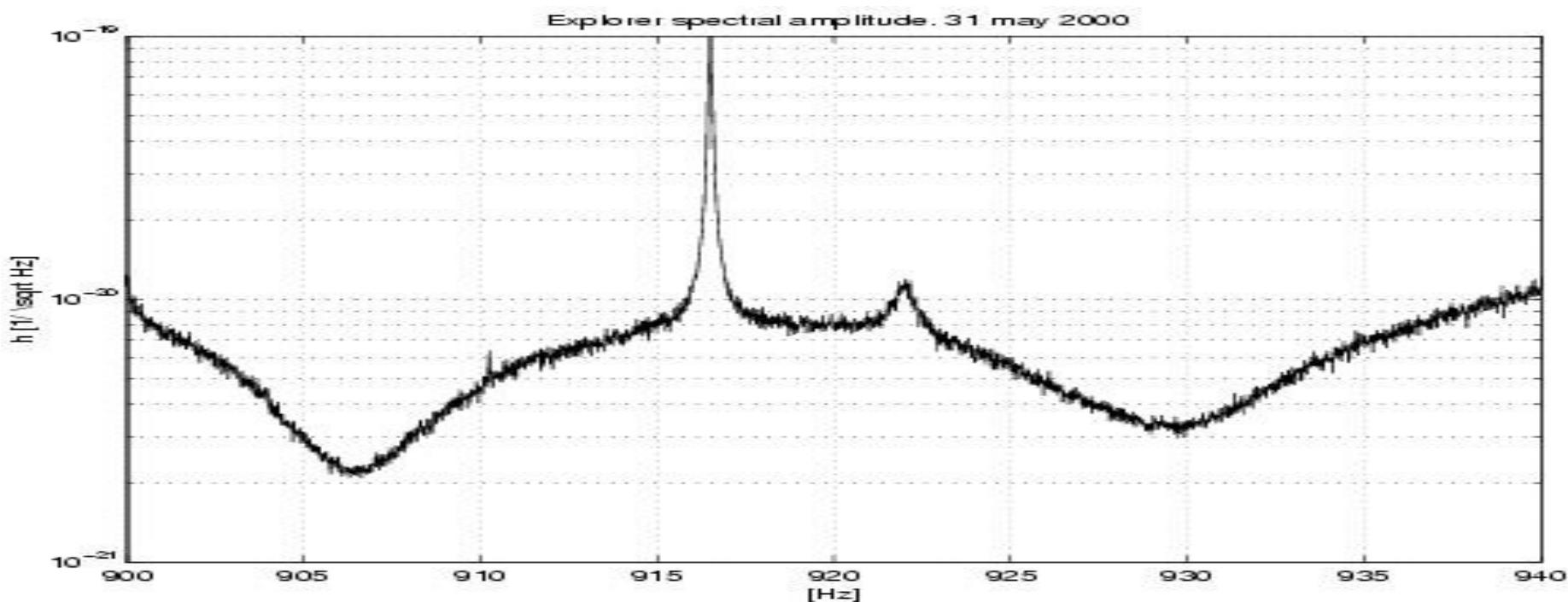


Virgo Data

Principle: Michelson laser interferometer with two Fabry-Perot cavities in the arms and power recycling.	
Arms length:	3 kilometers
Vacuum tube:	120 cm diameter, fully welded (no seal)
Vacuum performance:	
Pressure after baking at 1500 C:	< 10 ⁻⁹ mbar
hydrocarbons partial pressure	10-13 mbar
Laser and injection system:	
Nd:YAG diode-pumped laser at 1064 nm	
Laser power:	20 W
Mode cleaner:	144 m long, 3 mirrors ring cavity
Finesse of mode cleaner cavity	1000
Frequency stability	< 10 ⁻⁴ Hz ^{1/2} at 1 kHz
Relative Intensity Noise (RIN)	< 10 ⁻⁷ at 1 kHz
Power Stored in Fabry-Perot cavity:	15 to 50 kW (depending of finesse)
Power Stored in recycling cavity:	1 kW
Optics	
Main mirror substrates	special "Virgo" grade fused silica absorption < 1 ppm/cm diameter 350mm
Reflective coatings	absorption < 1ppm
Anti reflection coatings	reflection < 0.1 %
Surface quality	1/200 rms
Surface microroughness	< 1 Angstrom
Suspensions	
Inverted pendulum pre-isolator plus 6 passive stages.	
Active control on first and last stage and on mirror.	
Seismic attenuation	10-11 at 10 Hz
rms mirror motion before locking:	<1 mm, < 0.1 mradian
Thermal stabilization	+/- 0.1 degree C
Frequency range	
	10 Hz to 10 kHz
Sensitivity	
	$h \sim 3 \cdot 10^{-21} \text{ Hz}^{-1/2}$ at 10 Hz
	$h \sim 3 \cdot 10^{-23} \text{ Hz}^{-1/2}$ at 1 kHz







General	Spectro	Reconstruct	Inspirals	Bursts	Pulsars	Noise
DQ	FOM	Infrastructure	Vacuum	Locking	Alignment	Injection
Detection	Suspensions	Environment	GC	DAQ	Process	

Running processes

- The color of each CI boxes is a summary of the status of the processes monitored by this CI
- Red=some processes stopped , Orange=some processes not reachable by CI , Yellow=some processes in initialization step , DarkGreen=some processes active but not golden
- The number below the CI name is the number of processes in the state corresponding to the box's color

Click on a box to see the list of processes run by the corresponding CI.

This page is mainly dedicated to on-call experts should take care of any color not green

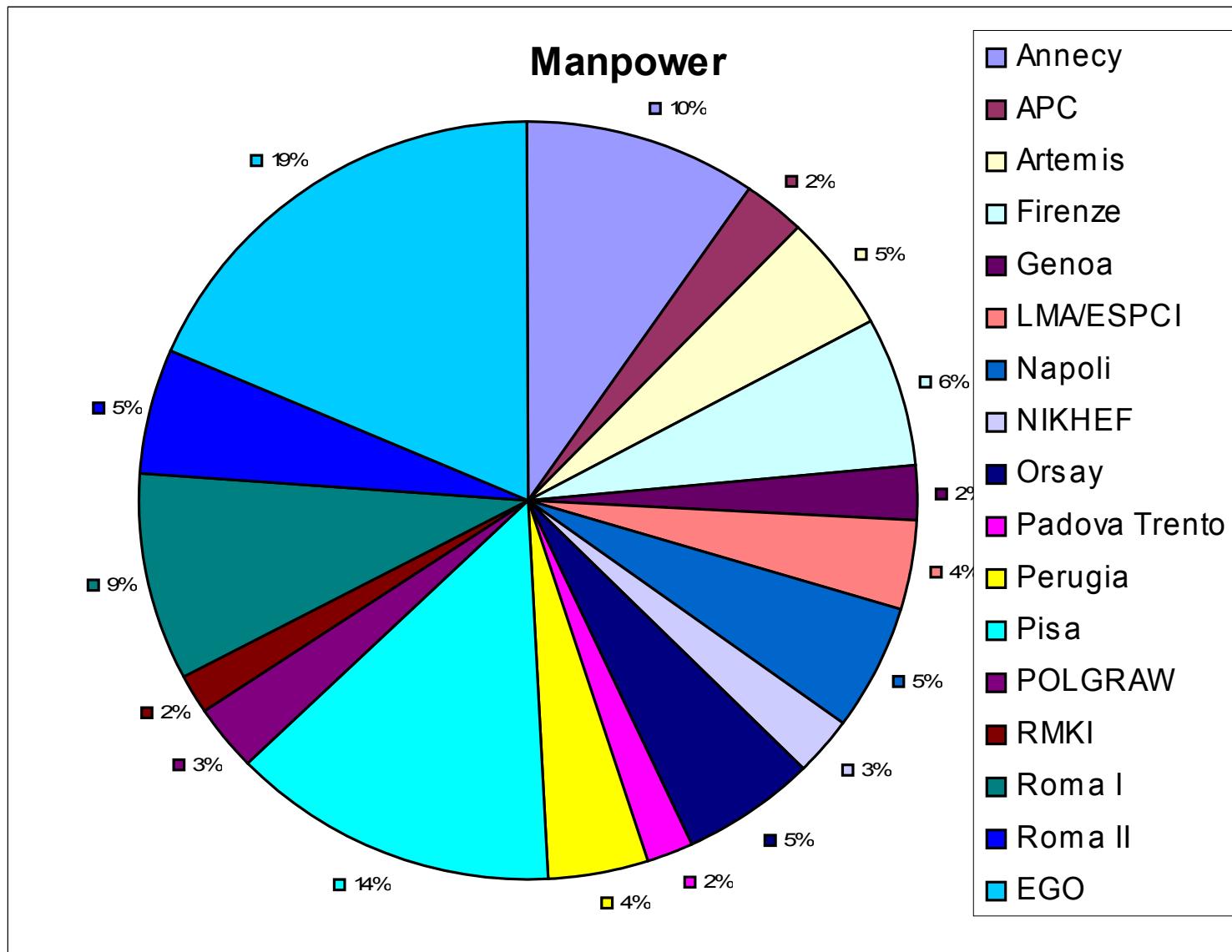
For operators: If a box is grey and indicates that the CI is inactive since xx minutes, it means that you must restart the CI in control room

Last update: Fri Oct 22 11:30:46 CEST 2010 - Last access: 22 października 2010 11:30:46

CIAlp Automation	ClBenchesE End Benches User Interface 4	ClBenches Central Benches User Interface 5	CICalibration Calibration 5	CHSYS ISYS Client
CITCS TCS Servers Client	ClEnvMon Environmental Monitoring Client	ClInfraMon Infrastructure Monitoring Client	ClITF ITF Control 1	CIDaqCollect DAQ Collect line 2
CIDaqData Data access line 1	ClDaqStorage DAQ Storage line	ClOP Online Processing line	CIDM Detector Monitoring line 2	CIDO Online DQ 1
CINoise Noise Monitoring line 2	CIWDF NoiseWDF Monitoring line 11	ClVeto Online Veto	ClCbcOnline inactive since 2637 mn	ClCbcHLV MBTA Trigger Production for H1, L1 and V1



h 9/2008



Total members: 256

POLGRAW contributions to VIRGO

- **Data analysis**

Participation in the all sky search for GWs emitted by rotating NS CW working group

Modeling populations and searching for GWs from coalescing NS and BHs including BH binaries with non-negligible spins

- **Detector characterization**

Search for periodic interferences in the data

- **Contribute to code simulating working and performance of the VIRGO detector**

- **Participation in data taking shifts**

Members of the POLGRAW

Kazimierz Borkowski, Centre for Astronomy, UMK, Toruń

Robert Budzyński, Instytute of Theoretical Physics, UW, Warsaw

Tomasz Bulik, Astronomical Observatory, UW, Warsaw

Piotr Jaranowski, Faculty of Physics, UB , Białystok

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Andrzej Pisarski, Faculty of Physics, UB, Białystok

Dorota Rosińska, Institute of Astronomy, UZ, Zielona Góra



Refereed Publications

Pia Astone, K. Borkowski, Piotr Jaranowski,
(Maciej Piętka), Andrzej Królak

Astone P. et al., 2002, Data analysis of gravitational-wave signals from spinning neutron stars. IV. An all-sky search, *Phys. Rev. D*, **65**, 042003.

Astone P. et al., 2003, All-sky upper limit for gravitational radiation from spinning neutron stars, *Class. Quantum Grav.*, **20**, S665–S676.

Astone P. et al., 2005, An all-sky search of EXPLORER data, *Class. Quantum Grav.*, **22**, S1243–S1254.

Astone P. et al., 2006, All-sky search of EXPLORER data: search for coincidences, *Class. Quantum Grav.*, **23**, S687–S692.

Astone P. et al., 2008, All-sky search of NAUTILUS data, *Class. Quantum Grav.*, **25**, 184012.

Astone P. et al., 2010?, Data analysis of gravitational-wave signals from spinning neutron stars. V. A narrow-band all-sky search, *Phys. Rev. D* (*in preparation*).